Assembling the earliest galaxies Pratika Dayal



Collaborators: James Dunlop, Umberto Maio, Benedetta Ciardi

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The tool: cosmological simulations



Ultraviolet Luminosity Functions

Dayal et al. 2013



The LBG UV spectral slopes



Negligible (no) contribution from PopIII stars even at $z\sim7$, 8 in galaxies detectable with JWST (HST).

Assembling high-z LBGs



Building up the major branch mass

Dayal et al., 2013



Progenitors of largest z~6 LBGs start assembling first with progenitors of smaller systems forming at progressively lower redshifts.

The major branch mass buildup: SF or mergers?



At any z, most major branch mass built up by star formation in the major branch with mergers contributing tiny amount to the total mass.

A common story: stochastic SF



Building up the major branch luminosity Dayal et al., 2013



Galaxies at the bright end gently build up their luminosity i.e. a positive luminosity evolution while galaxies at the faint end undergo a positive and negative luminosity evolution as they brighten and fade

The evolving UV LF: density + luminosity evolution



• Evolution of the bright end solely due to an increase in the luminosity

• Evolution of the faint end due to an evolution in both the luminosity and number density

The story of high-z galaxy assembly:

Rate of stellar mass assembly increases with decreasing redshift from z~12 to z~6.

Majority (~90%) of stellar mass of z~6 LBGs assembled by star formation in the major branch, with only 10% brought in by mergers.

W UV LF evolution depends on luminosity range probed:

- genuine physical luminosity evolution at bright end
- faint end is a mix of positive and negative luminosity and density evolution.